#### Local and gauge invariant observables in gravity

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### The need for local observables

Consider a Classical or a Quantum Field Theory on an n-dim. spacetime M.

- ▶ In QFT,  $\langle \hat{\phi}(x) \hat{\phi}(y) \rangle$  is singular for some pairs of (x, y).
- ▶ In classical FT,  $\{\phi(x), \phi(y)\}$  is singular for some pairs of (x, y).
- Instead, use smearing

$$\phi(\tilde{\alpha}) = \int_{M} \phi(x) \alpha(x) \,\mathrm{d}\tilde{x}$$

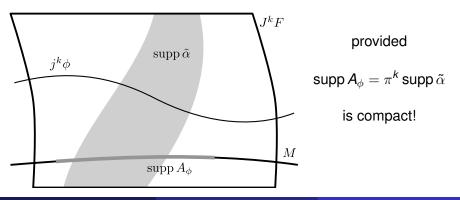
so that  $\langle \hat{\phi}(\tilde{\alpha}) \hat{\phi}(\tilde{\beta}) \rangle$  and  $\{ \phi(\tilde{\alpha}), \phi(\tilde{\beta}) \}$  are always finite, provided

- $\tilde{\alpha}, \tilde{\beta}$  are **smooth** *n*-forms on *M*,
- $\tilde{\alpha}$ ,  $\tilde{\beta}$  have **compact** supports.
- Smoothness diffuses singularities.
  Compactness ensures convergence of all integrals.
- Support of a functional: supp  $\phi(\tilde{\alpha}) = \operatorname{supp} \tilde{\alpha} \subset M$ .

#### Local observables

- Field  $\phi$  is a section of some bundle  $\pi \colon F \to M$  ( $\pi^k \colon J^k F \to M$ ).
- Local observables may be non-linear and depend on derivatives (jets). An *n*-form α̃ = α(x, φ(x), ∂φ(x), ···) dx̃ on J<sup>k</sup>F

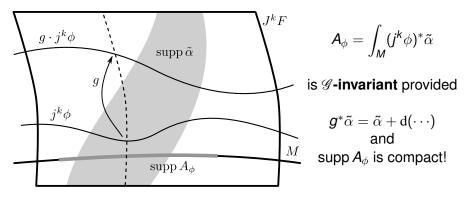
defines a local observable 
$$A_{\phi} = \int_{M} (j^{k} \phi)^{*} \tilde{\alpha},$$



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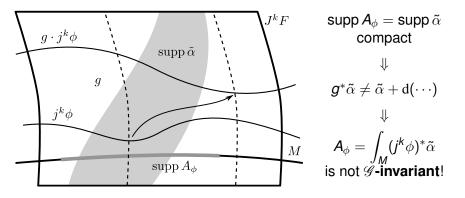
### Local observables in gauge theory (no gravity)

- ► Let 𝔅 be the group of gauge transformations.
- Gauge transformations  $g \in \mathscr{G}$  act on  $J^k F$  (hence  $j^k \phi \mapsto g \cdot j^k \phi$ ).
- ▶ No gravity:  $\mathscr{G}$  fixes the fibers of  $\pi^k : J^k F \to M$ .



### No (such) local observables in gravity

- Gravity is General Relativity (GR),  $F = S^2 T^* M$ ,  $\mathscr{G} = \text{Diff}(M)$ .
- Diffeomorphisms do not fix the fibers of π<sup>k</sup>: J<sup>k</sup>F → M. In fact, diffeomorphisms act transitively on these fibers.
- M is never compact, as needed by global hyperbolicity.



## Relaxing locality: an explicit example

• Take dim M = 4. Write the **dual Weyl tensor** as

$$\overset{*}{W}_{ab}{}^{cd} = W_{abc'd'}\varepsilon^{c'd'cd} = \varepsilon_{aba'b'}W^{a'b'cd}.$$

Make use of curvature scalars (Komar-Bergmann 1960-61)

$$b^{1} = W_{ab}{}^{cd}W_{cd}{}^{ab}, \qquad b^{3} = W_{ab}{}^{cd}W_{cd}{}^{ef}W_{ef}{}^{ab}, b^{2} = W_{ab}{}^{cd}\overset{*}{W}_{cd}{}^{ab}, \qquad b^{4} = W_{ab}{}^{cd}W_{cd}{}^{ef}\overset{*}{W}_{ef}{}^{ab}.$$

- ▶ Let  $\varphi$  be a **generic** metric (det  $|\partial b^i / \partial x^j| \neq 0$ ) and let  $\beta = (b^1[\varphi](x), b^2[\varphi](x), b^3[\varphi](x), b^4[\varphi](x))$  for some  $x \in M$ .
- Take a: ℝ<sup>4</sup> → ℝ, with sufficiently small compact support containing β, let α̃ = a(b) db<sup>1</sup> ∧ db<sup>2</sup> ∧ db<sup>3</sup> ∧ db<sup>4</sup> on J<sup>k≥2</sup>F

and 
$$A_{\phi} = \int_{M} (j^{k} \phi)^{*} \tilde{\alpha}.$$

A<sub>φ</sub> is well-defined on a Diff-invariant neighborhood U ∋ φ among all metrics φ such that R[φ]<sub>ab</sub> = 0. A<sub>φ</sub> is Diff-invariant.

## Differential invariants of fields (algebra)

- ▶ In any gauge theory, the group  $\mathscr{G}$  of gauge trans. acts on  $J^k F$ .
- **Differential invariants**: scalar  $\mathscr{G}$ -invariant functions on  $J^k F$ .
- **Theorem** (Lie-Tresse 1890s, Kruglikov-Lychagin 2011):
  - (generically) all differential invariants (all  $k < \infty$ ) are generated by
  - a finite number of invariants and
  - a finite number of differential operators satisfying
  - a finitely generated set of differential identities.
- Examples
  - Non-gauge theory: every function on  $J^k F$ .
  - Yang-Mills theory: invariant polynomials of curvature d<sub>A</sub>A.
  - Gravity: curvature scalars, built from Riemann R,  $\nabla R$ ,  $\nabla \nabla R$ , ...
- Gauge invariant observables: let α̃ = a(b<sup>1</sup>,..., b<sup>m</sup>) db<sup>1</sup> ∧···∧ db<sup>n</sup>, for some a: ℝ<sup>m</sup> → ℝ and differential invariants b<sup>i</sup>, i = 1,..., m ≥ n,

then 
$$A_{\phi} = \int_{M} (j^{k} \phi)^{*} \tilde{\alpha}$$
 is well-defined and gauge invariant,

provided supp  $[(j^k \phi)^* \tilde{\alpha}]$  is compact.

# Moduli spaces of fields (geometry)

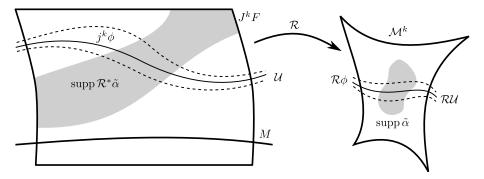
- In any gauge theory, the group  $\mathscr{G}$  of gauge trans. acts on  $J^k F$ .
- **Moduli space**: quotient space  $\mathcal{M}^k = (J^k F \setminus \Sigma^k) / \mathscr{G}$  ( $\Sigma^k$  is singular).
- Differential invariants are coordinates, separating points, on  $\mathcal{M}^k$ .
- Denote by R: J<sup>k</sup>F → M<sup>k</sup> the quotient map. Two (generic) field configurations φ and φ are gauge equivalent iff the images of Rφ(M) and Rφ(M) coincide as submanifolds of M<sup>k</sup> (for high k).
- Differential identities among differential invariants define a PDE *E<sup>k</sup>* on *n*-dimensional submanifolds of *M<sup>k</sup>*, identifying submanifolds like *R*φ(*M*).
- ► Finite generation means that there exists a k' such that all M<sup>k</sup> and E<sup>k</sup> (k > k') can be recovered from M<sup>k'</sup> and E<sup>k'</sup>.
- Choose compactly supported *n*-form α̃ on M<sup>k</sup> and U such that φ ∈ U implies Rφ(M) ∩ supp α̃ is compact. Then U is G-invariant,

$$A_{\phi} = \int_{M} (j^{k} \phi)^{*} \mathcal{R}^{*} \tilde{\alpha}$$
 is well-defined and gauge invariant,

and the  $A_{\phi}$  separate  $\mathscr{G}$ -orbits in  $\mathcal{U}$ .

## New notion of local and gauge invariant observables

- A<sub>φ</sub> may only be defined on an open subset U ⊂ S of (covariant) phase space. Local charts!
- ►  $A_{\phi} = \int_{M} (j^{k} \phi)^{*} \tilde{\alpha}$ , with  $j^{k} \phi(M) \cap \text{supp } \tilde{\alpha}$  compact for every  $\phi \in U$ .
- $A_{\phi}$  is gauge invariant if  $\tilde{\alpha} = \mathcal{R}^* \tilde{\beta}$  of some *n*-form  $\tilde{\beta}$  on  $\mathcal{M}^k$ .



▶ **NB:** Two metrics  $\phi$  and  $\psi$  are Diff-equivalent iff  $\mathcal{R}\phi = \mathcal{R}\psi$  in  $\mathcal{M}^k$ .

### Poisson brackets

- Poisson brackets of gauge invariant observables are well-defined intrinsically, but require care to compute.
- ► 1st possibility. Use hyperbolic gauge fixing to obtain Peierls bracket E(-, -),

$$\{A_{\phi}, B_{\phi}\} = E(A'_{\phi}, B'_{\phi}) = \int_{M \times M} A'_{\phi}(x) \cdot E_{\phi}(x, y) \cdot B'_{\phi}(y),$$

where  $A_{\phi+t\psi} = A_{\phi} + tA'_{\phi}(\psi) + O(t^2)$  and  $A'_{\phi}(\psi) = \int_M A'_{\phi}(x) \cdot \psi(x)$ , with  $A'_{\phi}(x)$  a compactly supported distribution.

- ► 2nd possibility. Use the reduced equation *E<sup>k</sup>* on *M<sup>k</sup>* and apply the Peierls formalism (write as hyperbolic PDE + constraints, linearize, compute Green functions).
- For gravity: 1st possibility is well understood. 2nd possibility not yet explored.

### Conclusion

- Local gauge invariant observables are important in both Classical (non-perturbative construction) and Quantum (perturbative or semi-classical renormalization) Field Theory.
- Usual restriction on "compact support" excludes gravitational gauge theories.
- Relaxing the support conditions opens the door to a large class of gauge invariant observables (even for gravitational theories), defined using **differential invariants** or **moduli spaces** of fields. They separate gauge orbits on open subsets of the phase space.
- > The **Peierls formalism** computes their Poisson brackets.
- Limitations:
  - Observables may not be globally defined on all of phase space.
  - Naive approach separates only generic phase space points (e.g., metrics without isometries).
  - Need to connect with operational description of observables.

Thank you for your attention!